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A modeling approach for water and sediment transport in Jezero crater on Mars based on new geomorphological evidence

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1. Introduction

Jezero crater – the investigation site of the current NASA's Mars 2020 mission – has been actively studied for almost 20 years. Since it was first described in detail in [[Fassett](#page-16-0) and Head, 2005], it was believed to once have contained a lake. Not only the geomorphological evidence – two inlets, entering the crater from the west and the north, the outlet valley, exiting the crater to the east and two delta deposits in the west and north accordingly to inlets – support this idea, but also the detection of clay minerals (first – from orbital data [[Poulet](#page-17-0) et al., 2005], [[Bibring](#page-16-0) et al., [2006](#page-16-0)], [[Mangold](#page-17-0) et al., 2007], [[Murchie](#page-17-0) et al., 2009]), which form under the presence of water. These arguments in favor of a paleolake in Jezero led to its selection as the target landing site for Mars 2020, one of the goals of which is to search for signs of possible ancient life. As at least episodic access to liquid water is almost universally accepted to be necessary for life (e.g. [\[Mottl](#page-17-0) et al., 2007]; [\[Jaumann](#page-17-0) et al., 2014]), constraining the fluvial activity in Jezero can contribute to its habitability assessment.

Jezero crater (diameter \sim 50 km) is located at the highland/lowland boundary on Mars [\[Tanaka](#page-17-0) et al., 2014] between the Nili Fossae region to the northwest, Isidis Planitia to the east and Syrtis Major to the southwest. The Nili Fossae region (Middle Noachian age, [\[Tanaka](#page-17-0) et al., [2014\]](#page-17-0)) is dominated by large grabens concentric to the Isidis impact basin, the grabens are considered to be the tectonic response to the basin formation ([\[Comer](#page-16-0) et al., 1985], [\[Wichman](#page-17-0) and Schultz, 1989], [[Goudge](#page-17-0) et al., 2015]). Isidis Planitia, which was formed in the Noachian as a result of an impact, was later resurfaced by volcanic and fluvial/glacial activities in the Late Noachian – Early Amazonian and by wind activity since the early Amazonian [[Ivanov](#page-17-0) et al., 2012]. The Syrtis Major volcanic province has Early Hesperian age [[Hiesinger](#page-17-0) and Head, [2004\]](#page-17-0). The Jezero crater itself is interpreted to be formed in the Noachian period: younger than the Isidis impact basin (~3.85–4.05 Ga

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[[Robbins](#page-17-0) et al., 2013]) and older than the first fluvial activity within the crater (based on the crater counts on the valleys of the Jezero, the fluvial system is estimated to be of ~3.8 Ga [[Fassett](#page-16-0) and Head, 2008], [[Goudge](#page-17-0) et al., [2015\]](#page-17-0)). The fluvial activity in the crater is constrained to the Noachian-Hesperian boundary, based on crater counting of the valleys [[Fassett](#page-16-0) and Head, 2008] and the western delta [\[Mangold](#page-17-0) et al., 2020].

Recent spectroscopic and geologic in situ observations by the Perseverance rover found igneous rocks altered by water [\[Farley](#page-16-0) et al., [2022\]](#page-16-0), [[Corpolongo](#page-16-0) et al., 2023], [Rice et al., [2023\]](#page-17-0) and even detected specific fluorescence-mineral associations consistent with many classes of organic molecules [[Sharma](#page-17-0) et al., 2023]. These findings support the idea that Jezero was a paleolake in the past. Several researchers tried to reconstruct, at least partly, the history of the paleolake using different approaches: 1) geomorphological analysis of the crater and valley discharge estimations [\[Fassett](#page-16-0) and Head, 2005], 2) analysis of the clayrich fluvial–lacustrine sediments of THEMIS and CRISM data [\[Ehlmann](#page-16-0) et al., [2008\]](#page-16-0), 3) analysis of the scroll bars and epsilon cross-bedding on the delta [[Schon](#page-17-0) et al., 2012], 4) a quantitative geomorphological study of the inlet valleys [\[Mangold](#page-17-0) et al., 2020], 5) modeling of the five waterrelated events (western valley carving, crater flooding by the western inlet, the eastern rim erosion, water outflow during the breaching and western delta deposition) [\[Salese](#page-17-0) et al., 2020], 6) photogeologic mapping ([[Stack](#page-17-0) et al., 2020], [[Jodhpurkar](#page-17-0) et al., 2024]), 7) study of the stratigraphic relations between the main western delta deposit, isolated mesas and crater floor units [\[Holm-Alwmark](#page-17-0) et al., 2021], 8) images of the delta outcrop taken by Perseverance [[Mangold](#page-17-0) et al., 2021], 9) hydrological models to constrain the climatic and hydrological conditions in Jezero [Horvath and [Andrews-Hanna,](#page-17-0) 2024] and 10) topographic detrending techniques to accentuate topographic features of the volcanic floor to study possible old locations of lakes and rivers in Jezero [Annex and [Ehlmann,](#page-16-0) 2024]. Moreover, several attempts were undertaken to model the outlet formation of Jezero – [Fassett and [Goudge,](#page-16-0) [2017\]](#page-16-0), [Holo and Kite, [2017](#page-17-0)], [Fassett and [Goudge,](#page-16-0) 2021].

In our research we were inspired by the study of [[Salese](#page-17-0) et al., 2020] and modeled all (ten) possible water-related processes using sediment and water transport models developed by [\[Kleinhans,](#page-17-0) 2005]. Moreover, we took into account the new data from the Perseverance rover [[Stack](#page-17-0) et al., [2024\]](#page-17-0), as well as careful measurements and our geomorphological observations based on the latest available remote sensing data to get more plausible input data to the model. The modeled events (processes) included: 1) western inlet (Neretva Vallis) valley carving, 2) northern inlet (Sava Vallis) valley carving, 3) crater flooding by only northern inlet and 4) by both northern and western inlets, 5) western rim breaching, 6) eastern rim breaching 7) water outflow from the crater, 8) outlet valley (Pliva Vallis) carving, 9) western delta deposition, 10) northern delta deposition (Fig. 1). Under "valley carving" we mean erosion of the valleys by flowing water; under "rim breaching" – erosion of the rim by ponding water, which overflows the edge of the rim. The goal of our study was to calculate the minimum timescales for each event and estimate the minimum volume of water provided/released during each event. The modeling results together with geomorphological observations allowed us to draw new conclusions about water activity in the crater and answer the questions: 1) which inlets were involved in the crater flooding; 2) when in the relative chronology of the crater's fluvial activity the northern and western deltas were deposited; and 3) whether there was enough water to flood the crater repeatedly and maintain open lake conditions for a long time.

Our study started with the definition of the model, mapping deltas and valleys, their morphometric measurements, as well as crater volume and diameter determination, and eastern rim breach and western rim breach volumes measurements (Section 2. Methods; [Section](#page-6-0) 3.1. Results of measurements). The geomorphological analysis ([Section](#page-8-0) 3.2. Geomorphological observations) added new information, which had to be taken into account during modeling. For example, we identified 2 incision events of the valleys; showed the importance of including the northern inlet to the model (which was not done in previous researches);

Fig. 1. Jezero's main features, studied in this paper (HRSC Mars Chart (HMC_13E10) DTM (50 m/pixel), overlayed by corresponding ortho-rectified image mosaics).

for the first time we also analyzed the western rim breach to include it in the model. After that we modeled ten water-related events, applying a variety of input parameters to study a range of possible scenarios ([Section](#page-9-0) 3.3. Results of modeling). We checked the uncertainties of results (coming from the uncertainties of the data, methods and the model itself) to set limits for our interpretation [\(Section](#page-10-0) 4.1. Understanding of the obtained results). After that we interpreted the obtained timescales ([Section](#page-12-0) 4.2. Relative comparison of timescales) and minimum available of water for each event [\(Section](#page-12-0) 4.3. Calculation of the minimum available water amount). We also considered the "fresh unfilled crater" case and showed that increasing the crater volume will not influence our interpretations [\(Section](#page-14-0) 4.4. Discussion, "Missing Sediments and the Fresh Crater Case"). Finally, we compared our interpretation with previous studies in [Section](#page-14-0) 4.4. Discussion, "Comparison With Previous Studies". Based on these studies, we achieved a better understanding on how the water-related (water transport and sediment transport) processes and events in the Jezero crater were interconnected, how they influenced each other and how they overlapped with each other.

2. Methods

2.1. Model description

We use flow discharge and sediment transport models for Martian conditions developed by Kleinhans [\[Kleinhans,](#page-17-0) 2005] which have been already applied many times to hydrological systems of Mars (for example, [\[Jaumann](#page-17-0) et al., 2010], [\[Hoke](#page-17-0) et al., 2011], [\[Mangold](#page-17-0) et al., [2012\]](#page-17-0), [\[Adeli](#page-16-0) et al., 2016]). Goal of the Kleinhans' models is to calculate water and sediment transport timescales under constant bank-full discharge. These are the conditions when most of the erosion occurs, because bank-full discharge can be assumed as the channel forming discharge for fluvial channels [[Leopold](#page-17-0) et al., 1964]). Bank-full conditions are conditions when water fills the whole channel to its banks but does not overtop them; discharge then is maximum [[Leopold](#page-17-0) et al., [1964\]](#page-17-0)). Studies on the frequency of bank-full discharges on Earth show different cycles for different regions – from half a year for many English rivers to 10 years in Pennsylvania [\[Harvey,](#page-17-0) 1969]; a recent study for the United States shows a 1.5-year cycle, although there are substantial variations around this numbers [\[Simon](#page-17-0) et al., 2004]. Another estimation is that channel-forming discharges on Earth occur for about 5 % of the fluvial active time [[Parker,](#page-17-0) 2005]. For Mars there are only some estimates on the intermittency factor for fluvial activity: for the Parana´ Valles region near the Noachian–Hesperian boundary it yields \sim 0.01–0.1 % during 10^5 – 10^6 yr ([Buhler](#page-16-0) et al., 2014). However, the question of the periodicity of the bank-full cycles on Mars as well as the longevity and periodicity of lake activity [\(Michalski](#page-17-0) et al., 2022) remains outstanding.

It is important to note that climate is not the only factor which determines bank-full conditions – river cross-section, channel slope, porosity of the substrate, grain size of the sediment and other geological and climatological factors also play a role [\[Kleinhans,](#page-17-0) 2005]. For example, the type of the water source (ground-water sapping/precipitation/snow melting/etc.) determines the response to intensive precipitation. As shown by [Laity and [Malin,](#page-17-0) 1985; [Howard](#page-17-0) et al., 1988; [Dunne,](#page-16-0) 1990], reaction of the sapping channel networks to precipitation is slower than that of the overland flow networks. Therefore, it is reasonable to assume that in this situation, a river fed by groundwater sapping would have a longer cycle of bank-full conditions than a precipitation-fed river. Moreover, river cross-section and channel slope also influence the bank-full cycle duration: it tends to increase as the river slope increases [\[Chau](#page-16-0) et al., 2021].

The Kleinhans models are based on water and sediment balances. For the water transport events they answer the question how long it takes for a certain amount of water to flow through a given channel (in bank-full conditions) and for the sediment transport events they can be used to compute how long it takes to transport a certain amount of sediment through a given channel (in bank-full conditions). Therefore, timescales are calculated by dividing the transported amount of the water by the water discharge and by dividing the transported amount of the sediment by the sediment transport rate. Climate conditions are not involved in the model; single water events could be separated by long durations. Therefore, timescales obtained here are minimum values. For example, deltas could have formed by the multiple episodes of water flow and sediment deposition. In our study we generalize the episodes of delta formation and consider the entire period of delta formation as a single delta-forming event in order to get an estimate of quantitative parameters.

A detailed description of the model is written in [Kleinhans,](#page-17-0) 2005. The formulas have been rewritten to show our step-by-step calculation approach (the desired parameter is always on the left side of the equation).

For the water transport timescale calculation, the channel size (width w , depth h , slope s), the volume of the transported water V_w and the median grain size D_{50} should be measured or estimated. Martian gravity g is known and taken as 3.72 m/s². First, the hydraulic radius $\rm R_h$ and the friction factor *f* (after Eq. 13 in [Kleinhans,](#page-17-0) 2005) should be calculated in order to estimate the friction of the channel floor and walls on the flow:

$$
R_h = \frac{w \cdot h}{w + 2h} \tag{1}
$$

$$
f = \frac{8}{\left(2.2 \cdot \left(\frac{h}{D_{50}}\right)^{-0.055} \cdot s^{-0.275}\right)^2}
$$
 (2)

Second, the Darcy-Weisbach equation [[Silberman](#page-17-0) et al., 1963] is used to calculate the depth- and width-averaged flow velocity:

$$
u = \sqrt{\frac{8 \cdot g \cdot R_h \cdot s}{f}} \tag{3}
$$

Third, channel width and depths together with flow velocity allow to calculate water discharge:

 $Q_w = w \cdot h \cdot u$ (4)

Finally, the timescale of the water transport event is obtained by dividing the transported water volume by the discharge:

$$
T_w = \frac{V_w}{Q_w} \tag{5}
$$

Sediments can be transported as bed load (along the stream bed by rolling or sliding) or as suspended load (which is uplifted in the flow) and it is possible to model both transport modes. *To calculate the timescale of sediment transport under suspended load conditions*, channel sizes (w, h, s) , the volume of the transported sediment (V_s) , median grain size (D_{50}), the sediment density ρ_s and the porosity λ should be measured or estimated. The Martian gravity $g(3.72 \text{ m/s}^2)$ and the fluid density ρ_w (for water – 1000 kg/m³) are known. First, the friction factor *f* is obtained using Eq. (2). Second, the relative submerged density – a dimensionless measure of sediments buoyancy – is calculated:

$$
R = \frac{\rho_s - \rho_w}{\rho_w} \tag{6}
$$

The third step is the calculation of the bed shear stress (Eq. (7)), which acts on the sediments covering the bed and initiates their motion. Following [Kleinhans,](#page-17-0) 2005, it is derived as the force of a block of water of unit length and width, with the water depth *h* over the channel bed and the slope *s*. In that case it is called the total shear stress of the flow because it is related to the total friction of the channel and includes both the bedform-related shear stress and grain-related shear stress. The nondimensionalized bed shear stress – the Shields parameter (Eq. (8)) – is calculated then from the bed shear stress:

$$
\tau_{total} = \rho_w \cdot g \cdot h \cdot \sin s \tag{7}
$$

$$
\theta = \frac{\tau}{(\rho_s - \rho_w) \cdot g \cdot D_{50}}\tag{8}
$$

The non-dimensional total transport rate (Eq. (9)) is taken as sediment transport predictor, which was postulated by [Engelund](#page-16-0) and Hansen, [1967](#page-16-0) and used in [Kleinhans,](#page-17-0) 2010. Subsequently, the Specific Volumetric Transport Rate (Eq. (10)) is calculated, which describes sediment transport per unit width per unit time $(m³$ per m width per second) and accounts for the pore space λ. After that, the Volumetric Transport Rate (Eq. (11)) can be obtained by multiplying by the channel width:

$$
\Phi_s = \frac{0.1}{f} \theta^{2.5} \tag{9}
$$

$$
Q_{s_spec} = \left(\frac{1}{1-\lambda}\right) \cdot \phi_s \cdot \sqrt{R \cdot g} \cdot D_{50}^{\frac{3}{2}} \tag{10}
$$

$$
Q_s = Q_{s_\text{spec}} \cdot w \tag{11}
$$

Finally, the timescale of sediment transport assuming suspended load $(Eq. (12))$ is calculated by dividing the volume of the transported sediment V_s by the volumetric transport rate:

$$
T_s = \frac{V_s}{Q_s} \tag{12}
$$

To calculate sediment transport timescale under bed load conditions, the friction factor, the relative submerged density, the hydraulic radius and the velocity are calculated according to formulas (2), (6), (1), (3). Then, the 90th percentile grain size D_{90} (and skin friction) can be calculated following the empirical relation, which was used in [Salese,](#page-17-0) [2020:](#page-17-0)

$$
D_{90} = 5 \cdot D_{50} \tag{13}
$$

Next, the grain friction factor (Eq. (14)) is taken as roughness predictor from the White-Colebrook function [[Silberman](#page-17-0) et al., 1963] (see Eq. 8 in [Kleinhans,](#page-17-0) 2005); the value D90 is chosen as hydraulic roughness length k_s (following [Kleinhans,](#page-17-0) 2005). The grain shear stress (Eq. [\(15\)\)](#page-3-0) is then calculated to represent the shear force on the grains which

is needed for the computation of the near-bed sediment transport (Eq. 17 in [Kleinhans,](#page-17-0) 2005):

$$
f_{gr} = \frac{8}{\left(5.74 \cdot \log\left(12.2 \cdot \frac{R}{D_{90}}\right)\right)^2}
$$
(14)

$$
\tau_{gr} = \frac{1}{8} \rho_w f_{gr} u^2 \tag{15}
$$

After that the grain-related non-dimensionalized grain shear stress – the grain-related Shields parameter (Eq. (16)) – is calculated (eq. 22 in [Kleinhans,](#page-17-0) 2005). It is needed to obtain the non-dimensional total transport rate (eq.17), which is also called the Meyer-Peter and Mueller predictor [\[Meyer-Peter](#page-17-0) and Mueller, 1948], [[Kleinhans](#page-17-0) and van Rijn, [2002\]](#page-17-0). Here the Shields criterion for incipient motion (θ_{cr}) is included (eq. 30 in [Kleinhans,](#page-17-0) 2005), which could be calculated using different empirical relations (see Section "4.3. Predictive Bed State Criteria and Diagram" in [Kleinhans,](#page-17-0) 2005). Here we use $\theta_{cr} = 0.03$ as in Kleinhans, [2010:](#page-17-0)

$$
\theta_{gr} = \frac{\tau_{gr}}{(\rho_s - \rho_w) \cdot g \cdot D_{50}}\tag{16}
$$

$$
\Phi_b = 8 \left(\theta_{gr} - \theta_{cr} \right)^{1.5} \tag{17}
$$

Last steps include the specific volumetric transport rate (Eq. (18)), the volumetric transport rate (Eq. (19)), and the timescale of sediment transport assuming bed load (Eq. (20)):

$$
Q_{b_{spec}} = \left(\frac{1}{1-\lambda}\right) \cdot \Phi_b \cdot \sqrt{R \cdot g} \cdot D_{50}^{\frac{3}{2}} \tag{18}
$$

$$
Q_b = Q_{b_spec} \cdot w \tag{19}
$$

$$
T_b = \frac{V_b}{Q_b} \tag{20}
$$

To decide which transport mode (suspended load or bed load) is more probable, the suspended/bed load transport ratio ("S/B ratio") is used: if it is far larger than one, then the system is suspension-dominated [[Kleinhans,](#page-17-0) 2005; [Salese](#page-17-0) et al., 2020]. [Kleinhans](#page-17-0) et al., 2010 decide to use five as a threshold. In our study we followed this approach as it also gives the widest possible range of timescales than using a threshold equal to one.

2.2. Methods of measurements

We use the Mars 2020 Science Investigation CTX DEM Mosaic (20 m/ pixel) and corresponding ortho-mosaic (5 m/pixel) for mapping the deltas and for subsequent volume measurements ([[dataset] [Calef,](#page-16-0) [2021\]](#page-16-0), [[Malin](#page-17-0) et al., 2007]); valley mapping and volume measurements were based on the HRSC Mars Chart (HMC_13E10) DTM (50 m/pixel) and the related ortho-rectified image mosaics (12.5 m/pixel) [\[Gwinner](#page-17-0) et al., [2016\]](#page-17-0). The HRSC Mars Chart (HMC) data was accessed via the map-based data portal at <https://maps.planet.fu-berlin.de> [\(Walter](#page-17-0) et al., [2018\)](#page-17-0) and converted to the IAU (International Astronomical Union) system with a sphere radius of 3,396,190 m. For the channel measurements we use both the Mars 2020 Science Investigation CTX DEM and Ortho-Mosaics and the Mars 2020 Terrain Relative Navigation HiRISE DTM Mosaic (1 m/pixel) with the corresponding Ortho-Mosaics (0.25 m/pixel) ([\[Fergason](#page-16-0) et al., 2020], [\[Malin](#page-17-0) et al., 2007]). The data were co-registered to HRSC and reprojected to a sinusoidal projection.

Before measuring, the deltas and the valleys were mapped at a scale of 1:5000. The valleys were mapped using two morphological features (Fig.SM1–5, Supplementary Material). We assume that they represent two distinct phases in valley development and are not expressions of different lithologies. The first is the initial old valley, which has no borders that are recognizable in image data. It was mapped based on slope and curvature rasters derived from the HRSC DTM. The second is the last incised valley, which was mapped on the HMC ortho-rectified image mosaic. The difference between the borders can be seen in [Fig.](#page-4-0) 2, [Fig.](#page-4-0) 3 and Fig. SM1–5 (Supplementary Material). In the upper section of the western inlet (starting from the valley head, to about 130 km to the east) it appears that the initial valley is broader than the last incised valley, whereas in the lower section (the easternmost 70 km) the visible borders of the initial and the last incised valleys almost coincide ([Fig.](#page-4-0) 3). This is explained by the fact that the lower section is covered by an olivine-bearing unit [[Mangold](#page-17-0) et al., 2020]. It blanketed the lower part of the initial valley and the interior and the rim of the Jezero crater [[Mangold](#page-17-0) et al., 2020; [Goudge](#page-17-0) et al., 2015]). As a result, any former fluvial and lacustrine deposits within this area were also buried by this unit. The lower section of the last incised valley was carved into the olivine-bearing unit after its emplacement. Therefore, the volume of the initial valley is underestimated in its lower section. Where the boundaries of the initial valley were not traceable in the Slope raster, we have drawn them in the same way as the valley boundaries of the last incision (based on the HMC mosaic).

To map the last incised valleys of inlets and outlet, first, certain and uncertain borders of the valleys were identified on the HMC orthomosaics as well as the flow traces. However, the borders of the northern valley are not continuous, therefore they were approximated in the most realistic way using information of the slope gradient direction derived from the HMC DTM [\(Fig.](#page-5-0) 4).

As a result, all three valleys – western (Neretva Vallis), northern (Sava Vallis) and outlet (Pliva Vallis) were mapped in 2 extents – initial valley and last incised valley extents (Fig. SM1–5, Supplementary Material).

Both deltas were mapped with 3 possible extents: minimum, medium and maximum ([Fig.](#page-5-0) 5 and [Fig.](#page-5-0) 6). The minimum extent for the western delta is based on the GIS version of the map of [[Stack](#page-17-0) et al., 2020]. Delta map units "blocky", "layered rough", "thickly layered", "thinly layered", "truncated curvilinear layered" were merged into one single polygon for the purpose of this study. For the medium extent the western borders remain the same while the eastern borders were set according to the nearest buttes. The maximum extent includes controversial areas to the north and to the west (near the crater rim) while in the south and in the east the extent is following the farthest buttes.

Mapping the extent of the northern delta is more challenging, because it is heavily eroded. For the minimum extent we followed the steep edge (although sometimes it is destroyed and only remnants are left) in the west and in the south (without buttes). In the east we followed the border from the map of [Jodhpurkar](#page-17-0) et al., 2024 and [Jodh](#page-17-0)[purkar](#page-17-0) and Bell, 2024, who mapped two geological units based on a change of surface properties along this border. The medium extent in general follows the borders of the minimum extent, but in the south the closest buttes are included. The maximum extent includes controversial areas in the west (closer to the eroded crater rim), in the north the border starts at the valley entrance to the crater, in the east we followed the borders from the map of [Goudge](#page-17-0) et al., 2013, and in the south the most distal buttes were included.

Measurements were conducted in ArcGIS. Width, depth, and slope of the channels were measured using longitudinal and cross-sectional profiles on the CTX DTM. [Salese](#page-17-0) et al., 2020 observed a 5 m-high terrace within the western valley at the entrance into the Jezero crater (see Fig. 2 in their paper) and used it as a bank-full channel depth of the western inlet. Following this methodology, we also found a channel terrace in the northern valley (with the average depth 4.6 m) and a channel terrace in the outlet valley (with the average depth 13 m) (see Fig. SM8–9, Supplementary Material); these values were taken as channel depths in the model. The water volume needed to fill the crater before breaching was determined as the crater volume up to the elevation of the first breaching terrace (recognized at the eastern rim where the outlet is exiting the crater – see [Fig.](#page-9-0) 8 and [Section](#page-8-0) 3.2. Geomorphological observations) using the standard Surface Volume Tool. It

Fig. 2. Segments of the last incised western valley (a), mapped on the HMC ortho-mosaic (b) and the initial valley (c), mapped on slope raster (d) and profile curvature raster, calculated from the HMC DTM. e) – Elevation profile through the initial valley and last incised valley (on HMC DTM) which shows that the boundaries of the initial valley are located at the convex bends of the valley profile (points A and D). Therefore, they are drawn along the area between the steep slope (yellow color) and a flat surface (green color) – Fig. 2c and d. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

Fig. 3. Initial western valley (mapped on slope and profile curvature rasters, calculated from the HMC DTM) and the last incised western valley (mapped on the HMC ortho-mosaic).

calculates the volume between the DTM and the horizontal plane of a certain elevation (in this case − 2243 m). Similarly, the minimum volume of water that outflowed during the eastern rim breaching was estimated. This amount of water was stored between the level of the 1st breaching terrace and the outlet floor at the eastern rim. The third type of measurements is the volume of the eroded and deposited sediments. The eroded volume of the western rim breach and the eastern rim breach (not only the total volume, but also the volumes eroded during each

Minimum extent

Medium extent

Maximum extent

Fig. 5. Western delta in 3 extents (description of borders in the text).

Minimum extent

Medium size

Maximum size

Fig. 6. Northern delta in 3 extents (description of borders in the text).

breaching phase) was estimated using the Surface Volume Tool and information about the elevations of each breaching terrace. Our method accounts not only for the changing distances between terraces for each breaching phase, but also for the narrowing of the rim ridge upward.

To measure volumes of the valleys, the mapped borders were transformed into a sequence of points with an interval of 50 m using the "Generate Points Along Lines" tool in ArcGIS. To each point the height information was added from the DTM using the "Add Surface Information" tool. After that, the "Natural Neighbor" tool was used to interpolate the surface which existed before the valley carving. Finally, the obtained surface was overlapped with the actual DTM of the valley in the "CutFill tool" to calculate the volume of the eroded rock. To crosscheck the results of this method, we roughly estimated the volume by multiplying the valley area by its mean depth. The results were in good agreement.

For estimating the volumes of the deltas two methods were used: 1)

minimum, medium and maximum volumes were calculated by multiplying the delta area by the mean delta thickness (40 m for the northern and 60 m for the western delta, measured on cross-section profiles); 2) the actual delta volume was measured using the CTX DTM. For this, a tilted plane was created using information about the delta slope (assuming that the surface under the delta has the same slope as the delta itself) – Fig. 7. This plane, lowered by the mean delta thickness (60 m), represents the underlying surface. Overlapping with the actual DTM in the "CutFill tool" gives the volume of the deposited sediments.

The measurement uncertainties were also assessed. The first measurement uncertainty arises from distortions associated with the map projection. To assess this uncertainty, we followed the methods of [Butcher](#page-16-0) et al., 2016 and [Butcher](#page-16-0) et al., 2020 by comparing the lengths of 1◦ meridional lines in different places over the study area, measured in planar and geodetic mode. The differences were minor $(-1, 9)$.

The uncertainty of the vertical measurements and thus the volume measurements directly depend on the internal vertical error of the DEM or DTM, which means the same thing on Mars. For the Mars 2020 Science Investigation CTX DEM Mosaic, which is used for most of the measurements with exception of valley lengths and volumes, the intrinsic (relative) vertical error originating from merging of the several stereo combinations into the final DTM mosaic is described separately for the western part and for the eastern part of the mosaic. The western part of Mosaic has relative vertical offsets of 3.8 m between the individual CTX DEMs [\[Fergason](#page-16-0) et al., 2020]; for the eastern part only an approximate qualitative assessment is available – vertical offsets are not greater than \sim 10 m here (dataset author Fred Calef, personal communication, February 6, 2024). For the HRSC Mars Chart DTM the overall internal consistency of the data can be described by the mean intersection error which summarizes the precision of all object points [[Gwinner](#page-17-0) et al., 2009]. For example, for the MC-11 quadrangle subset the mean intersection error for all points in multi-orbit DEM is not more than 8.9 m [\[Gwinner](#page-17-0) et al., 2016]. Approximately the same error is expected for the MC-13 quadrangle subset which was used in this study. The Mars 2020 Terrain Relative Navigation HiRISE DTM Mosaic has a median vertical offset to the CTX DTM mosaic of the Lander Vision System elevation map equal to ~ 0.4 m.

The uncertainties due to the human factor and manual digitalization of the features by the choice of exact points to draw the borders of objects are usually insignificant in morphometric studies and do not exceed 1 % [\[Butcher](#page-16-0) et al., 2016].

3. Results

3.1. Results of measurements

The measured key parameters are shown in [Table](#page-7-0) 1. They can be classified into three groups: 1) characteristics of the channels – width, depth, slope – to calculate discharge and sediment transport rate; 2) volume of the crater – to estimate water volume needed to fill the crater before the eastern rim breaching; 3) delta volumes, valley volumes and carved rim volumes – to estimate the amount of transported/deposited sediments. The obtained values are consistent with those obtained in previous studies ([\[Stack](#page-17-0) et al., 2020], [[Salese](#page-17-0) et al., 2020], [[Mangold](#page-17-0) et al., [2020\]](#page-17-0), [[Mangold](#page-17-0) et al., 2021], [Gazetteer of [Planetary](#page-17-0) Nomenclature, International [Astronomical](#page-17-0) Union (IAU), 2024], [\[Ehlmann](#page-16-0) et al., [2008\]](#page-16-0), [[Schon](#page-17-0) et al., 2012]). This shows that the developed measurement methods are reliable and that the values obtained for features which were not studied before are plausible (northern inlet measurements, western rim breach, outlet measurements).

The western delta and the western valley are larger than correspondingly the northern delta and the northern valley. The amount of material eroded from the valleys during the last incision $(8 \text{ km}^3 \text{ for the}$ western valley and 2 km^3 for the northern) is approximately the same as deposited in deltas (1.44–4.58 $km³$ for the western delta and 1.58–3.73 $km³$ for the northern). That means that, with high probability, the deltas were deposited during the last incision of the valleys. [Goudge](#page-17-0) et al., [2018](#page-17-0) reported about the incision on top of the western delta which they mapped as Jezero Western Delta (JWD) valley (see Fig. 2 in their paper). This would imply that the last cutting phase was after the delta was deposited. However, this does not contradict our observations and interpretations, because JWD valley (width 500–600 m, depth 30–50 m – see Fig. 7 in [Goudge](#page-17-0) et al., 2018) is smaller than our mapped western last incised valley (mean width \sim 1200 m, depth \sim 140 m) and one valley can be nested in another. Moreover, channel which was carved during the last cutting phase after the delta was deposited [\(Fig.](#page-4-0) 2 in [Goudge](#page-17-0) et al., [2018\)](#page-17-0) would have smaller size within the last incised valley mapped in our study, as there it would cut the more competent bedrock to the west of the crater rim in contrast to less competent alluvial material within the delta itself. The deposition of the transported sediments from the initial valleys will be discussed in [Section](#page-14-0) 4.4. Discussion. The lengths of the main tributaries of the western and northern valleys are comparable – 210 km and 160 km, accordingly. However, the western inlet (\sim 3000 m width and \sim 200 m depth) is significantly more incised than the northern inlet (\sim 2000 m width and \sim 100 m). This shows that the western inlet was formed by more extensive erosion, hinting at a longer period of fluvial activity. It is worth mentioning that valleys today might be mantled with aeolian deposits and mass wasting material

Fig. 7. a) Profiles of the delta surface and obtained underlying surface; b) Plan view of the underlying surface.

Table 1

Morphometry of the Jezero crater. * For the "tilted plane method" explanation see [Section](#page-3-0) 2.2. Methods of measurements. References in the table: Gazetteer of Planetary Nomenclature, International Astronomical Union (IAU)(2024), Ehlmann et al. (2008), Fassett and Head (2005), Goudge et al. (2015); Mangold et al. (2020), Mangold et al. (2021), Salese et al. (2020), Schon et al. (2012),

In previous studies

Table 1 (*continued*)

Measured feature In this In this

(*continued on next page*)

Table 1 (*continued*)

and were deeper in the past; therefore, depths of valleys and channels and their volumes obtained in this study are minimum values.

In the outlet also an initial valley and a last incised valley were found (Fig. SM5, Supplementary Material), which could be interpreted as 2 erosion events or 2 phases of erosion within the outlet formation. The distinctly visible part of its length is shorter (60 km) than the inlet valleys, however, taking into account sudden interruption of the valley due to possible post-formational erosion of its downstream part, it could be longer (*>*80 km) and flow directly to the dichotomy border which is located nearby (to the east). Another possibility is that the flow could spread out into different directions. There is at least one bifurcated channel, which is visible in CTX imagery. Therefore, there could also be more channels, smaller and shallower, continuously branching into more streams, creating a sheetwash towards the lowlands at the dichotomy boundary. The third possible reason for the sudden valley disappearance may have been the infiltration of the water into a highly porous rock or underground drainage systems with sinkholes and caves, although this explanation is speculative since there are no sink hole detections in the vicinity of Jezero so far. Finally, the channel erosion could be limited due to transition to deposition which happened in the presence of the hypothetical standing body of water in the Isidis basin.

Concerning channel metrics, we found that mean channel depths and widths in the western inlet and the northern inlet are almost the same (for the western $-$ ~190 m width and \sim 5 m depth; for the northern – \sim 196 m width and \sim 5 m depth). However, the channel slope is steeper for the northern inlet channel (~ 0.0132) than for the western (~0.0095), resulting in higher velocity for nearly the same cross-section assuming bank-full flow. This in turn implies higher discharge. In the outlet the mean channel depth $({\sim}13 \text{ m})$ and width $({\sim}400 \text{ m})$ are greater than in both inlets, while the slope (-0.0125) is comparable to the northern inlet.

We also determined the volumes of the western and eastern rim breaches. The amount of carved rim material is 1 km^3 and 3.1 km^3 , respectively. We did not conduct measurements for the northern rim breach because we assumed that the entire northern segment of the rim was already eroded by the time of fluvial activity and had approximately

the same morphology which we see now. It could be eroded by the ancient aeolian activity or, for example, by possible coastal abrasion when Jezero was submerged under Arabia ocean [\(Baum](#page-16-0) et al., 2021). This assumption is plausible and confirmed by our results (see [Section](#page-12-0) [4.3.](#page-12-0) Calculation of the minimum available water amount). In addition, the absence of the northern segment of the rim makes the reconstruction of the northern rim breach impossible.

The diameter of the crater, the crater volume below −2243 m (the flood stage before eastern rim breaching) and the minimum volume of water outflow during breaching were re-measured and agree with the values obtained by previous researchers (see [Table](#page-7-0) 1).

3.2. Geomorphological observations

3.2.1. Terraces in the inlets and outlet

In [[Salese](#page-17-0) et al., 2020] it was shown that there are 3 breaching terraces at elevations − 2243 m, − 2280 m and − 2310 m at the outlet rim which allowed them to hypothesize 3 different phases within the breaching event (Fig. 2R in [\[Salese](#page-17-0) et al., 2020]). We also analyzed a profile along the eastern rim and found these breaching terraces [\(Fig.](#page-9-0) 8) to be in agreement with Salese's hypothesis.

Our geomorphological analysis of the northern inlet reveals terraces at approximately the same height as the terraces at the eastern rim ([Fig.](#page-9-0) 9). The elevation difference between recognized terraces in the northern inlet and terraces at the eastern rim does not exceed 10 m, which lies in between the vertical errors of the DTM (measurement uncertainties are discussed in [Section](#page-3-0) 2.2. Methods of measurements). These observations indicate that the northern inlet was active during the eastern rim breaching and, therefore, was involved in the flooding phase which later led to the creation of the outlet at the eastern rim. Another point is that these terraces lie outside the last incised valley, indicating that they were formed before the last incision event. The terraces on the right river bank (to the west from the northern inlet) are apparently accumulative, while terraces on the left river bank (to the east from the northern inlet) are due to an erosive origin – as indicated by the valley bend on this site.

The western inlet, in contrast, has no terraces at the breaching levels of the eastern (outlet) rim (-2243 m, -2280 m and -2310 m). This means that it either did not participate in crater flooding, or its terraces were eroded, or that the water flow at the time of the eastern rim breaching was above these levels. Later (in [Section](#page-12-0) 4.3. Calculation of the minimum available water amount) we show that, even if the western inlet was involved in the first crater flooding, the western rim was more likely not carved down to these levels (− 2243 m, − 2280 m and − 2310 m) at the time of the eastern rim breaching.

However, the western rim exposes one distinct terrace at the elevation − 2085 m and two small terrace fragments (or morphologic steps) of uncertain origin at the elevations − 1800 m and − 1990 m ([Fig.](#page-10-0) 10). According to topography to the west from the crater, during breaching of the western rim the water could not rise above the level of approximately − 2085 m, because otherwise it would overflow the river banks and flow to the south and to the north around the crater. In this case it would be impossible for water to pond in front of the western rim and breach it. Therefore, we modeled the western rim breaching as one phase breaching between levels of − 2342 m and − 2085 m.

3.2.2. Western and northern deltas

Both deltas lie below the − 2360 m isoline (Fig. SM7, Supplementary Material). Assuming that they were not much higher before they were eroded, this indicates that they were formed when the water level in the lake was *<*− 2360 m. Deposition therefore started during the last breaching phase (which happened between the 3rd eastern rim terrace at a level of − 2310 m and the outlet floor at the eastern rim at a level of − 2410 m). A similar conclusion was published in [\[Salese](#page-17-0) et al., 2020] for the western delta. However, on the DTM in this paper the western delta lies below − 2410 m, suggesting delta deposition occurred after

Fig. 8. Profile through the eastern rim showing the three breaching terraces, the maximum width and depth of the breached channels and the volume of material eroded (in red) during each breaching phase. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

Fig. 9. Terraces at the northern inlet lie approximately at the same heights as the breaching terraces at the eastern rim, indicating that the northern inlet was active during the eastern rim breaching, therefore, also participated in the crater flooding before breaching.

eastern rim erosion, water outflow and subsequent lake level decrease to approximately − 2410 m (the height of the outlet floor at the eastern rim). Nevertheless, it is just a small difference compared to our measurements. The main result from our observation is that carving of the last incised valleys coincided with the deposition of the deltas, and these 2 events happened not earlier than during the last breaching phase. These findings are in agreement with [Mangold](#page-17-0) et al., 2024, who found out that delta deposition occurred during a rapid decrease of the water level.

We do not exclude the possibility of the existence of a paleo-delta in the past, which might have been later covered and eroded. This would explain the fate of the 43 $km³$ eroded from initial valleys (see also [Section](#page-14-0) 4.4. Discussion).

3.3. Results of modeling

The obtained timescales of water-related events are presented in [Table](#page-11-0) 2. The modeling results in detail for each water-related event can be found in the Supplementary Material (Tables SM1–14).

For modeling, the Martian gravity was taken as 3.72 m/s^2 , water density as 1000 kg/m³, sediment density as 3500 kg/m³ as assumed for Martian basaltic rock by [Kleinhans,](#page-17-0) 2005 and [Salese](#page-17-0) et al., 2020, sediment porosity as 0.2 as assumed by [Kleinhans](#page-17-0) et al., 2010 and [Salese](#page-17-0) et al., [2020,](#page-17-0) and Shields criterion for incipient motion as 0.03 as in [Kleinhans,](#page-17-0) 2005 and [Salese](#page-17-0) et al., 2020. The recent study by [Beyssac](#page-16-0) et al., [2024](#page-16-0) showed that the olivine-rich boulders in the upper fan and margin unit at Jezero crater are olivine cumulates, which texture and

Fig. 10. Distinct terrace at the elevation − 2085 m and two small terrace fragments of uncertain origin at the elevations − 1800 m and –1990 m. Breaching was modeled up to terrace at − 2085 m (see explanation in text). "View Point" is placed to show the perspective view (Fig. SM6, Supplementary Material).

composition align with those of terrestrial dunite. Dunite is a highdensity rock (dunite density in Martian meteorite is 3460 kg/m³ – [Coulson](#page-16-0) et al., 2007) which is consistent with the value of sediment density assumed in this study.

3.3.1. Valley carving and delta deposition

The timescales for eroding the northern valley, the western valley and the outlet valley were calculated for both the initial valley volume and last incised valley volume. To assess the timescales of the initial valley incisions, we used the same discharge as calculated for the channel measured within the last incised valley, assuming discharge was approximately similar during both incision periods. In [Section](#page-14-0) 4.4. Discussion we show that even decreasing and increasing the dimensions of the initial valley channel by a factor of 2 will not influence our final interpretation.

We modeled a range of scenarios of Median Grain Size (D50) for the valley carving and delta deposition. Following [Salese](#page-17-0) et al., 2020, we assume the more alluvial environment for Jezero's western and northern inlets, therefore we used D50 from 0.005 m to 0.014 m, which correspond to pebble. The timescales for the intermediate D50 values – 0.008 m, 0.01 m, 0.012 m – were also calculated. For the outlet valley there is no in-situ information about its geology, therefore we also assume 0.005 m as a minimum possible grain size for its carving. The maximum D50 for the outlet valley is taken as 0.1 m, because outlet formation follows eastern rim breaching and could be considered as a breaching-related process and $D50 = 0.1$ m was already assumed for the eastern rim breaching by [Salese](#page-17-0) et al., 2020.

3.3.2. Western and Eastern rim breaching

The breaching events of the crater rim were modeled as both catastrophic breaching and long-term erosion under constant flow scenarios. For the catastrophic breaching scenario we assume a channel width in the rim equal to the distance between the rim remnants to the north and to the south at the corresponding breaching terrace height. For example, for the 1st breaching phase of the eastern rim it is 7550 m, see [Fig.](#page-9-0) 8. This follows the approach by Salese et al. [\(2020\)](#page-17-0). For the channel depth, we modeled the range of values starting from a minimum of 5 m to a maximum possible depth for each breaching phase. The maximum possible depth is equal to the distance between the corresponding terraces; for example, for the 1st breaching phase of the eastern rim the maximum possible depth of the channel is 37 m, see [Fig.](#page-9-0) 8. Intermediate values between 5 m and 190 m were included in the model with steps of 5 m. As there is no in-situ information about rim geology, we followed [Salese](#page-17-0) et al., 2020 assumption of $D50 = 0.1$ m (cobbles), although in Fassett and [Goudge](#page-16-0) (2021) they found that at $D50 = 0.1$ m the outlet

could hardly continue to erode due to the absence of the needed positive feedback of drainage and incision. Nevertheless, there is no better estimate and therefore it can be used as an upper limit of the grain size for the rim breaching. We also added the second scenario ($D50 = 0.02$ m, pebbles) to check the sensitivity of the model to the grain size in the case of rim breaching.

For the long-term erosion under constant flow scenario, we assume uninterrupted long-term flow in a channel of approximately the same dimensions as the present channel in the corresponding last incised valley (for the western rim of about 190 m width and 5 m depth and for the eastern rim of about 400 m width and 13 m depth). The slopes of the channel for each breaching phase were reconstructed using slope values of the nearest rim remnants to the north and to the south from the breach. We also consider median grain sizes of 0.1 m and 0.02 m for modeling the long-term erosion under constant flow.

3.3.3. Flooding of the crater and water outflow during breaching

Flooding the crater up to the 1st breaching terrace (− 2243 m) was modeled by only the northern inlet and by both the northern and western inlet. We did not model flooding by only the western inlet because, according to the geomorphological evidence, the northern inlet had to participate in the crater flooding before the eastern rim breaching. For the western inlet there is no such strong evidence (see [Section](#page-8-0) [3.2.](#page-8-0) Geomorphological observations).

As water outflow is connected with the eastern rim erosion, we included both the median grain sizes of 0.1 m and 0.02 m in the model (the same we used for the eastern rim erosion modeling). For the channel widths and depth, we followed the methodology already described for the western and eastern rim breaching (see above). We also corrected the water outflow for water coming from both the western and northern inlets, although the effect of this correction is negligible (difference within several hours).

4. Interpretation & discussion

4.1. Understanding of the obtained results

Before interpreting the results, their application should be understood. In this section we call up model limitations and model assumptions to understand the meaning of the obtained results – minimum timescales and minimum water amount for each water-related event.

The model has been already validated by previous researchers using knowledge about timescales of the catastrophically formed fan in southern Iceland [\[Duller](#page-16-0) et al., 2008; [Duller](#page-16-0) et al., 2014] and two recent river-dominated artificially originated terrestrial fan deltas: the Wax

Table 2

Timescales of water-related events.

Water-related event			Timescale (Earth years)
Northern Valley Carving		Initial valley	55-496
		Last Incised Valley	15-136
Western Valley Carving		Initial valley	407-3053
		Last Incised Valley	98-739
Outlet Valley Carving		Initial valley	$1 - 15$
		Last Incised Valley	$0.4 - 6.2$
Western Delta Deposition		Minimum extent (measured on DTM)	12.8-96.3
		Minimum extent ("tilted plane" method)	17.3-129.5
		Medium extent	24.7-185.3
		Maximum extent	54.9-412
Northern Delta Deposition		Minimum extent (measured on DTM)	$7.4 - 67.1$
		Minimum extent ("tilted plane" method)	11.7-106
		Medium extent	13.2-119.4
		Maximum extent	27.7-234.2
Western Rim Erosion	Catastrophic	$D50 = 0.1$	$0.0014 - 3.7$
		$D50 = 0.02$	$0.00034 - 2$
	Long-term	$D50 = 0.1$	38.5
		$D50 = 0.02$	20
Eastern Rim Erosion	Catastrophic	All breaching phases, $D50 = 0.1$	$0.016 - 2.2$
		All breaching phases, $D50 = 0.02$	0.00397-3.9
	Long-term	All breaching phases, $D50 = 0.1$	$4.6 - 7.2$
		All breaching phases, $D50 = 0.02$	9.3
Flooding the crater up to 1st breaching terrace		With northern inlet	$6.41 - 6.8$
		With northern and western inlets	$3.19 - 3.4$
Water outflow during breaching	Catastrophic	All breaching phases, $D50 = 0.1$	0.0031-0.1169
		All breaching phases, $D50 = 0.02$	0.0033-0.1285
	Long-term	All breaching phases, $D50 = 0.1$	0.35
		All breaching phases, $D50 = 0.02$	0.39

Lake Delta (Louisiana, USA) and the Lake Constance delta (Austria) [[Salese](#page-17-0) et al., 2020]. While for the catastrophic event of a fan in southern Iceland the timescale obtained by the model of 2–17 h almost coincided with the real timescale of 6–10 h, the results for 2 recent fan deltas (the Wax Lake Delta and the Lake Constance delta) show an error by a factor of 2: the Wax Lake Delta modeled timescale is 15 years with a real time of 37 year and the Lake Constance delta modeled timescale is 50 years with a real time of 115 years. These errors may originate from the uncertainty about bank-full conditions. Appling the rule of "channelforming discharges on Earth occur roughly for about 5% of the fluvial activity time" [[Parker,](#page-17-0) 2005] (already mentioned in [Section](#page-1-0) 2.1 Model description) would not help to remove the difference between the real

and modeled time (for the Wax Lake Delta, 15 years / 5% = 300 years \neq 37 years; for the Lake Constance delta, 50 years / 5% = 1000 years \neq 115 years). Periods of intensive flooding, when rivers can even overflow the banks, also play an important role in sediment transport and decrease the timescales obtained by the model and corrected for the "5% bank-full conditions time"-rule. This apparently explains the case with the Wax Lake Delta, whose fluvial system has undergone the high flood years of 1973–1975 with abnormally high sediment flux. A similar situation is present in the Lake Constance delta fluvial system which has extreme seasonal deviations from the average ("During the peak of the thaw period in the Alps, water flow can be 10 times greater, and the amount of suspended load can increase by more than a factor of 20 times" [\[Salese](#page-17-0) et al., 2020]). This demonstrates the importance of the type of the water source, in this case snow melting, for water and sediment transport (ground-water sapping and precipitation influence is already described in [Section](#page-1-0) 2.1. Model description). The uncertainty arising from the periodicity of the bank-full conditions, seasonal changes and also climatic changes is large on Earth, and on Mars even larger. Therefore, it can be said with confidence that timescales obtained by this model are not comparable with the real geological timescales of fluvial processes, but only show a model time of these processes based on water and sediment balances.

For the Jezero we do not know the periodicity of the bank-full conditions, and also the presence/absence and duration of dry periods with no fluvial activity are not known exactly. [Salese](#page-17-0) et al., 2020 noticed that the longer the real time of formation, the more important is the role of intermittency: "if the lake formed in a catastrophic event, then intermittency is not as relevant, but if it took centuries, then it is likely that discharge fluctuated over seasonal or longer time periods". For example, for the rim erosion timescales that means, if the breaching was catastrophic, the real timescales are close to those calculated by the model. However, if breaching was by long term gradual erosion, modeled timescales would differ from the real timescales. Gradual filling with water followed by overflowing water over the rim, and again filling the resulting void, again overflowing water, etc., and also accompanied by dry periods or periods of reduced water activity with lower discharges – all these factors will result in high intermittency and increase the real timescale of the water-related events in comparison with the modeled one. Moreover, initially evaporation and water infiltration into the underground are also not included in the model. In out study, they were assessed afterwards [\(Section](#page-12-0) 4.3 "Calculation of the minimum available water amount). In addition, the model also is quite sensitive to some input parameters, such as slope and grain size, as shown by [Salese](#page-17-0) et al., [2020.](#page-17-0)

Finally, the equilibrium erosion/bank-full flow model can underestimate the discharge of highly dynamic events like rim breaching and subsequent outlet valley carving. It is reasonable to expect that in the beginning of the valley carving the discharge at the outlet must be significantly higher than at the bank-full level, considering that initially there was no outlet channel. However, the best estimate of discharge that can be obtained within the model for the rim breaching is the bankfull discharge, which is based on the distance between the breaching terraces (see [Section](#page-9-0) 3.3. Results of modeling, Western and Eastern Rim Breaching). This approach was also used by [Salese](#page-17-0) et al., 2020.

Taking all the above considerations into account, timescales obtained by the model are not absolute (geological) timescales. However, they provide relative comparisons under the assumption that fluvial processes in Jezero were equally exposed to climatic factors and responded equally to their changes. This assumption is plausible as the northern inlet and western inlets were likely active at a comparable time (as shown in Scenarios 2 and 3 in [Fig.](#page-10-0) 10 in [Jodhpurkar](#page-17-0) et al., 2024). Therefore, the periodicity of the bank-full conditions in the western and northern inlets was the same; potential dry periods stopped activity in both inlets simultaneously and interrupted fluvial processes with the same periods of intermittency. Comparison of the timescales constraints the relative duration of the water-related events, their potential overlapping and interconnection. Another application of the results is calculation of the minimum available amount of water as minimum provided water by inlets and minimum outflow of water through the outlet. These can be compared with each other (to draw conclusions about the balance between water input and water output) as well as with Jezero's basin volume (to check if the northern inlet could alone flood the crater). Moreover, the water amount comparison also provides a cross-check for the conclusions made during the relative comparison of the timescales. Finally, our approach provides an estimate of the total water amount Jezero crater has experienced (see next section - 4.2. Relative comparison of timescales).

4.2. Relative comparison of timescales

4.2.1. Western rim erosion and crater flooding

Here we consider a possible case when the western inlet was participating in the first crater flooding together with the northern inlet and carved the western rim on its way to the crater. Comparison of the timescales of the western rim erosion with the timescales of flooding the crater up to the 1st breaching terrace leads to different conclusions depending on the different scenarios. For a catastrophic scenario, the western rim erosion happened as fast as the flooding of the crater or faster. It means, that, when the breaching was already completed, the crater was either still being flooded or had been completely filled. For the long-term scenario it is vice versa: the western rim erosion takes much longer than flooding the crater. In this case the crater was already flooded but the western rim erosion was not still finished. It also means that eastern rim breaching could start before the western rim erosion was finished. However, there is still a possibility that the flooding of the crater started only with the western inlet, and the northern inlet joined later; the opposite situation – the beginning of flooding by the northern inlet, and the joining of the western later – is also feasible. In these cases, an estimation of the timescales is difficult and relative duration of the western rim erosion and crater flooding is unknown.

Relative comparison of timescales of the western rim erosion and crater flooding favors neither catastrophic nor long-term scenario. However, the calculation of the minimum water amount needed for western rim erosion will allow us to establish a priority scenario (Section 4.3. Calculation of the minimum available water amount).

4.2.2. Eastern rim breaching and water outflow during breaching

The timescales for the water outflow during breaching are obtained for that amount of water (236 km^3) which was stored in the crater between levels of the 1st breaching terrace (− 2243 m) and the outlet floor at the rim (− 2410 m). Comparison of these timescales with the timescales for the eastern rim erosion (see [Table](#page-11-0) 2) shows that in most of the modeled cases in both long-term and catastrophic scenarios the eastern rim erosion lasted much longer than the water outflow. That means that 236 $km³$ of water has not been enough to carve the breach. However, it is possible to find overlapping timescales for both events in the catastrophic scenario: for $D50 = 0.1$ the resulting (overlapping) timescales are be 0.016–0.1169 Earth years (5.84–42.67 Earth days); for D50 = 0.02 the resulting (overlapping) timescales are 0.00397–0.1285 Earth years (1.45–46.9 Earth days). These timescales are comparable with previously obtained timescales for the outlet-forming flood in Jezero (9 Earth days 20 Earth hours – see Movie S4 in Fassett and [Goudge,](#page-16-0) 2021). Hence in a catastrophic scenario there is a possibility that both events – rim erosion and water outflow – happened simultaneously or, in other words, the eastern rim could have been eroded by only the 236 km^3 of water. For a long-term scenario such an overlap does not appear and in this case, a significant water supply from the inlets would be needed to finish eastern outlet rim carving.

4.2.3. Outlet valley carving and Eastern rim breaching

Comparison of the timescales obtained for the whole outlet valley carving (initial and last incised valleys together: 1.4–21.2 Earth years)

with the timescales of the eastern rim erosion in *a long-term scenario* $(4.6-7.23$ Earth years for the grain size $D50 = 0.1$ and 9.32 Earth years for the grain size $D50 = 0.02$) shows two possibilities:

- 1) for the case of 0.005 *<* D50 *<* 0.02 in the outlet valley, the timescales of the whole outlet valley carving (1.4–5.1 Earth years) are shorter than timescales for rim erosion for $D50 = 0.02$ (9.32 Earth years) (Table SM3). This means that the outlet carving should have been already finished while the rim erosion was still ongoing, which is impossible;
- 2) for $D50 = 0.1$ in the outlet valley, the timescales of the whole outlet valley carving (21.2 Earth years) are higher than the timescales for the rim erosion for $D50 = 0.1$ (4.6–7.23 Earth years). In this case rim erosion finished while outlet valley continued to carve, meaning open-basin lake conditions, which seems plausible.

In the catastrophic scenario, only in the case when breaching channel had depth 5 m and only if $D50 \le 0.014$ m for the outlet valley, the timescale of rim erosion (3.8 Earth years) is larger than timescales for the outlet valley carving (1.4–3.7 Earth years). This again means longer rim erosion than outlet carving, which is impossible. In all other cases the outlet valley carving would last longer than rim erosion, which points to open-basin lakes condition. To sum up, most of the modeled cases, in both long-term and catastrophic scenarios, support the idea of open-basin lake conditions after eastern rim breaching.

4.2.4. Northern last incised valley carving and Northern delta deposition

Timescales for the northern last incised valley carving (15–136 Earth years) coincide with timescales of the northern delta – in the medium extent 13.2–119.4 Earth years, in minimum extent 11.7–106 Earth years and in maximum extent 27.7–234.2 Earth years. This demonstrates hydrological connectivity and consistency between these events – last valley incision and delta deposition.

4.2.5. Western last incised valley carving and Western delta deposition

Timescales for the western last incised valley carving (98–739 Earth years) overlap with the timescales obtained for the deposition of the western delta – in the maximum extent 54.9–412 Earth years, in the medium extent 24.7–185.3 Earth years, in the minimum extent 17.3–129.5. This also confirms the hydrological connectivity between the last incision of the valley and delta deposition.

4.3. Calculation of the minimum available water amount

The minimum (because it is bank-full conditions) water volume provided during initial valley carving and the last incised valley carving is calculated by multiplying the corresponding timescales by the discharges [\(Table](#page-13-0) 3). For each event, the minimum volume of provided/ released water has a range of an order of magnitude for each event. This is explained by the range of possible scenarios we used in the model. Discharges for the last incised valley are calculated using metrics of the preserved channels within the valley (see Eq. [\(4\)](#page-2-0)). As it is not possible to determine the discharge of the initial valleys, the same discharge as for the last incised valley was assumed (in [Section](#page-14-0) 4.4 Discussion we show that increasing and decreasing of the initial valley channel sizes by a factor of 2 will not influence our interpretation). The obtained minimum volume of provided water is a multiple of the volume of Jezero's basin (See [Table](#page-13-0) 3, Columns 5th and 6th). This implies the possibility of a long-term period of fluvial activity and open-lake conditions. Moreover, it also shows that the northern inlet could alone fill the crater before breaching.

The comparison of the minimum amount of water discharged during eastern rim breaching (236 km³ was stored between the levels of the 1st breaching terrace and the outlet floor at the rim) with the water released through the outlet valley $(1000-14,800 \text{ km}^3)$ shows that the outlet valley could not have been carved only by the water stored in the crater.

Table 3

Minimum volume of water provided by inlets and released by outlet. Bank-full discharge (column 3) is calculated for the channel measured within the last incised valley.

Therefore, Jezero had more water input from the inlets and consequently must have been an open-basin lake during or after eastern rim breaching. This includes the possibility of a partly-open basin, which remains closed while water is accumulating behind the weir but becomes opened as soon as water overflows. This overflow can happen as catastrophic flooding (with the discharge much higher than bank-full), or as longer overspill (with approximately bank-full discharge), or, what is most likely, the combination of both. Therefore, the real value of discharge and therefore timescale lie between these 2 extreme cases – catastrophic breaching and long-term erosion. The timescales of the most of the modeled cases for the outlet valley carving and eastern rim erosion also support the idea of the open-basin lake (see [Section](#page-12-0) 4.2. Relative comparison of timescales, "*Outlet Valley Carving and Eastern Rim Erosion*"). Moreover, comparison of this amount of water of 236 km³ with the minimum amount of water needed to carve the eastern rim breach shows the same: in most scenarios more water is needed than just 236 km^3 (see Table 4). Only in the case of the catastrophic scenario with $D50 = 0.02$ and channel depth = 5 m this amount of water of 236 km³ would be sufficient to carve the breach, as only 194 km^3 would be

Table 4

Minimum amount of water needed for rim erosion (see also detailed Calculations 1–3 in the Supplementary Materials).

needed (see Table 4).

The minimum amount of water provided by inlets (38100–295,000 $km³$) is higher than the minimum amount of water released out of the lake (1000-14,800 $km³$). The possible explanation for this is that some water flowed through the outlet in non-bank-full conditions and parts could have been evaporated or/and drained into the underground. Taking the Jezero paleolake area as approximately 1900 km^2 , the thickness of water layer which was infiltrated vary from 19 km (37,100 km³/ 1900 km²) to 147 km (280,200 km³ / 1900 km²). According to [Shadab](#page-17-0) et al., 2024, infiltration capacity f_c for Noachian era for basaltic crust is 10^{-6} m/s (32 [m/year]). The infiltration of the abovementioned layer of water would last from 593 years $(19 \text{ [km]} / 32 \text{ [m/year]})$ to 4593 years (147 [km] / 32 [m/year]). These timescales of infiltration within the Jezero crater (593–4593 Earth years) are comparable with the timescales of studied events (for example, for the carving of the total western valley it is 505–3792 Earth years, see [Table](#page-11-0) 2). At the same time, evaporation could also be an explanation for the substantial difference between the water volume provided by inlets and the water volume released by the outlet. However, our estimations even with the high evaporation rates (0.1–1 m/year for Eberswalde crater for present conditions – [Irwin](#page-17-0) et al., 2015) give timescale of 190,000 –1,470,000 Earth years, which shows that water loss by evaporation is less significant than by infiltration.

Comparison of the minimum water volume needed for the western rim erosion (Table 4) with the water volume needed to fill the crater before breaching (446 km^3) shows that in most cases the crater flooding finished faster than western rim breaching. This is only possible in the long-term scenario (see [Section](#page-12-0) 4.2. Relative comparison of timescales). This comparison could also explain why there are no terraces in the western inlet which would lie at the same heights as terraces at the eastern outlet rim. In most of the modeled cases the western rim would not have been carved down to these levels when the eastern rim breaching started to happen. At the same time this conclusion confirms our assumption, that the northern rim was already eroded by the time of the northern inlet activity (see [Section](#page-6-0) 3.1. Results of measurements). If not, there would have been the same situation as in the western inlet and

we would not find any terraces at the same levels as the breaching terraces in the outlet on the eastern rim.

If the first crater flooding was a continuous process and was not interrupted by dry periods, enough water to flood the crater and cause the eastern rim breaching would be provided already during the first stages of the initial valley carving. Therefore, most of the volume of the valleys was eroded already after the eastern rim was breached.

4.4. Discussion

4.4.1. Missing sediments and the fresh crater case

If sediments from the last incision of the valleys were deposited in the northern and western deltas ([Sections](#page-6-0) 3.1. Results of measurements and [Section](#page-9-0) 3.3 Results of modeling), the question arises: Where are the \sim 43 km³ of sediments eroded from the initial valleys and the western rim breach? One possible explanation could be an interruption of fluvial activity and complete drying of the Jezero lake before the resumption of activity. During such an intermittency extensive wind erosion could occur, blowing away the missing sediments. The significant exhumation of the whole region and missing material around Jezero is also reported by [Bamber](#page-16-0) et al., 2022.

Another possibility is that basalt emplacement could bury the older sediments transported from the initial valleys and western rim breach. The rough estimation of the volume of the last visible basalt layer, which underlies the present delta ([[Russell](#page-17-0) et al., 2023], [\[Paige](#page-17-0) et al., 2024]) gives a value of \sim 11 km³ derived by multiplication of the volcanic unit maximum thickness of \sim 30 m [\[Schon](#page-17-0) et al., 2012] by its area of \sim 350 km². There could have been more basalt layers underneath. In the extreme case, sediments from the initial valleys were eroded and transported to the crater floor when the crater was still fresh without any infilling. To estimate the volume of the fresh crater we integrated a function describing the fresh crater profile from the center point to the rim, which was obtained using the empirical formulas from [Kleinhans](#page-17-0) et al., [2010](#page-17-0) and Garvin and [Frawley,](#page-17-0) 2003 (see also Supplementary Material, Calculations 4–6). Calculation of the volume of the fresh crater is based on its diameter. The diameter measurement yield uncertain results because the crater rim is partly eroded and it is difficult to find the crest line on the crater rim. In our study we obtained the value of 50 km for diameter (the same as in [[Stack](#page-17-0) et al., 2020], [\[Salese](#page-17-0) et al., [2020\]](#page-17-0)) using the 3 highest points on the rim to reconstruct the rim crest. However, also values of 45 km [\[Mangold](#page-17-0) et al., 2021] and 47.52 km [Gazetteer of Planetary [Nomenclature,](#page-17-0) International Astronomical Union [\(IAU\),](#page-17-0) 2024] were obtained by previous researchers. We calculated crater volumes for these diameters to show the sensitivity of this method to the diameter (Table 5): increasing the diameter by 5 km (\sim 11 %, from 45 km to 50 km) will increase the volume by 473 km³ (\sim 31 %, from 1549 to 2022 km 3) which makes a considerable difference. Table 5 shows that \sim 43 km³ of sediments missing from the initial valleys and western rim breach can easily be a part of the crater infilling which yields approximately 1100–1580 $km³$ (calculated as the difference between the fresh and the present crater volume).

Another question arises in connection with the fresh crater case: could the northern inlet alone flood the crater during the initial valley carving? For that, it had to fill the 2022 km^3 of the fresh crater with water, which is still less than the $3600-33,000$ km³ of water provided during the initial valley incision [\(Table](#page-13-0) 3). However, the timescales for

the northern initial valley carving were estimated based on the assumption that the initial valley contained the same channel with the same dimensions as the last incised valley do [\(Section](#page-9-0) 3.3. Results of modeling, *"Valleys Carving and Delta Depositions"*). But even if the real channel in the initial valley was 2 times larger or smaller than assumed in this study, the northern inlet alone could still flood the fresh crater completely (Supplementary Material, Calculation 7). In addition, the water volume calculated here is still a minimum value, as we consider only bank-full conditions. In non-bank-full conditions, even more water would be provided, leaving no doubt about the possibility of flooding the fresh crater by the northern inlet alone.

4.4.2. Comparison with previous studies

The evolution of the Jezero basin started with the flooding of the crater. In our study we showed that it was not done only by the western inlet, as previously thought [\(Salese](#page-17-0) et al., 2020), but either by the northern inlet alone or by the northern and western inlets together. The evidence for this is terraces in the northern inlet at the same heights as breaching terraces at the eastern rim breach. For the first time we also considered the western rim erosion which had to happen during the crater flooding. Most of the modeled cases show that the western rim erosion lasts longer than the crater flooding. We also showed that the northern rim was already eroded by the time of the fluvial activity, as indicated by the morphological comparison with the western rim and calculations of the minimum provided water (see [Section](#page-12-0) 4.3. Calculation of the minimum available water amount).

The second stage of evolution is breaching of the eastern rim in at least 3 episodes which were found by [Salese](#page-17-0) et al., 2020. We add to this that at that point, when the eastern rim erosion started, the western rim breaching was not yet completed (in most of the modeled cases, *>*95 %). We also found out that the amount of water stored in the crater between the levels of the 1st breaching terrace and the outlet floor at the eastern rim was not enough to erode the eastern rim breach and outlet valley. This implies that Jezero was an open-basin lake during or after the eastern rim breaching. These results are in agreement with previous studies [[Fassett](#page-16-0) and Head, 2005], [[Schon](#page-17-0) et al., 2012], [[Goudge](#page-17-0) et al., [2015\]](#page-17-0), etc. At the end of the eastern rim breaching stage, most of the last valley incision occurred and most of the volume of the northern and western deltas was deposited. The explanation for this is that both northern and western delta lie a little bit higher than the outlet floor at the eastern rim, implying that they started to deposit material already during the 3rd phase of the eastern rim breaching.

Finally, the lake level drops as soon as the water balance is negative. After a complete stop of the water supply it dries out/seeps out and disappears. In our study we also introduced the possibility of several events of drying and flooding, because such an intermittency in fluvial activity could explain the absence of the \sim 43 km³ of sediments carved from the initial valleys and the western rim. For this, the lake must have dried out at least once before the deposition of the present deltas, allowing wind erosion to blow away the sediments or deposition of basalts to cover them.

There is still a discussion if the eastern rim breaching as well as the western rim breaching, which was never studied in detail before and still raises questions (for example, the western inlet formation mechanism and its relationship with rim erosion is unknown – [\[Bamber](#page-16-0) et al., [2022\]](#page-16-0)), formed in catastrophic events or by long carving with multiple overflow events. In [Salese](#page-17-0) et al., 2020 there is no final answer to this question: "…it had at least three catastrophic collapses or it overflowed for a long (unknown) time such that it carved out the breach". Other researchers tend to have different points of view. The maturity of the outlet channel, according to [Schon](#page-17-0) et al., 2012, suggests the breaching in a non-catastrophic manner. The 1D model of Fassett and [Goudge,](#page-16-0) 2017 implied slow fluvial erosion (over tens of thousands of years), while 2D model (in BASEMENT) was consistent with catastrophic carving of the outlet under a narrow set of assumptions. Holo and Kite's, [2017](#page-17-0) results show that the observed incised depth of the outlet can be obtained

within \sim 100 years after the onset of sediment transport. [Goudge](#page-17-0) et al., [2019](#page-17-0) studied 24 basins (including Jezero) and found a correlation between total drained volume and Martian paleolake geometry. In the Jezero case this correlation supports that the outlet incision was the result of a single lake overflow flood, although they also considered multiple flooding events and incision by longer-term overflow. Finally, the model of Fassett and [Goudge,](#page-16-0) 2021 suggest that for large lakes like Jezero a single catastrophic flood is sufficient for outlet incision. Each model has its limitations – hydrodynamic and sediment transport assumptions, boundary conditions uncertainties, accuracy of the data, etc. In addition, the mechanism of the crater rim breaching on Mars – overtopping, piping or induced slumping [Fassett and [Goudge,](#page-16-0) 2021] – is also unknown. In our study we assumed an overtopping mechanism and modeled both catastrophic and long-term scenarios although for the long-term erosion still without possible intermittency between the short overflowing events. We found that the long-term scenario is the most probable for the western rim breaching, which lasted longer than the crater flooding. And for the eastern rim, the results of modeling do not prioritize any one scenario but show that in both the long-term and catastrophic scenarios breaching lasted much longer than the outflow of water stored in the crater before breaching.

5. Conclusions

In our study:

- 1) The water-related processes in the Jezero crater were modeled, including events which have not previously been fully or extensively studied: western rim erosion, flooding of the crater by only the northern inlet and by both the northern and the western inlet, northern delta deposition.
- 2) Other processes which were studied before by other researchers western valley carving, western delta deposition, eastern rim erosion and water outflow during breaching – were remodeled using a wider range of scenarios and our new measurements. For valley carving and delta deposition we incorporated findings from the Perseverance rover; for rim breaching we modeled both catastrophic and longterm scenarios.
- 3) Detailed geomorphological observations allowed us to identify: a) two stages of valley carving each of the inlets and the outlet; b) the presence of the terraces in the northern inlet (which strongly coincide with the height of the breaching terraces at the eastern outlet rim); c) the absence of such terraces in the western inlet, and d) the height position of the deltas relative to the outlet floor at the eastern rim.
- 4) Detailed mapping of the two stages of valley carving (initial valley and last incised valley) as well as delta mapping in 3 extents (minimum, medium and maximum) provided us the base for the geomorphic measurements.
- 5) Uncertainties of the used model are also carefully investigated to identify the limitations and put the interpretation on a stable basis.
- 6) The obtained timescales of the model constrain the available water for each water-related event within the Jezero crater.
- 7) The location of the missing sediments from the initial valleys and western rim breach was discussed with the possibility of the fresh crater flooding.
- 8) Our results and interpretations support and improve previous studies.

The conclusions of our study are presented in Table 6 (the most important conclusions are in bold). We show that the history, order and relative duration of the water-related processes in the Jezero crater could be more complicated than previously thought. We proved that the northern inlet was involved in the first crater flooding and subsequent eastern rim breaching. The inlets as well as the outlet had at least 2 incision events, and deltas were deposited during the last incision event.

Table 6

Conclusions and explanations (the most important conclusions are in bold).

(*continued on next page*)

Table 6 (*continued*)

Finally, each inlet, even separately, could provide enough water to flood the crater multiple times and maintain an open-basin lake within the crater. Our findings, although do not completely reconstruct the past of the Jezero paleolake in geological times, but still shed light on the fluvial and sediment transport processes and condition of the water history in Jezero.

CRediT authorship contribution statement

Anastasiia Ovchinnikova: Writing – original draft, Visualization, Methodology, Investigation, Conceptualization. **Ralf Jaumann:** Writing – original draft, Supervision, Investigation, Conceptualization. **Sebastian H.G. Walter:** Writing – review & editing, Software. **Christoph Gross:** Writing – review & editing, Validation. **Wilhelm Zuschneid:** Writing – review & editing, Validation. **Frank Postberg:** Writing – review & editing, Resources, Project administration.

Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

Data availability

The CTX and HiRISE data are available through the Planetary Data System Imaging Node at <https://pds-imaging.jpl.nasa.gov/>. The HRSC terrain models and image data are available at the data archives of ESA and NASA and in the map-based data portal at [https://maps.planet.](https://maps.planet.fu-berlin.de) [fu-berlin.de](https://maps.planet.fu-berlin.de) [\(Walter](#page-17-0) et al., 2018).The dataset with the shape-files of the mapped valleys and deltas in the Jezero crater is available at <https://doi.org/10.17169/refubium-45033> ([Ovchinnikova](#page-17-0) and Jau[mann,](#page-17-0) 2024).

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Appendix A. Supplementary data

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